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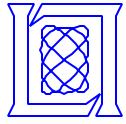
# Spectral Calibration of the EO-1 Advanced Land Imager\*

Jeffrey A. Mendenhall

Advanced Space Systems and Concepts  
Group

MIT/Lincoln Laboratory





# Outline

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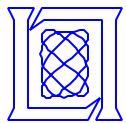
- **Introduction**
- **Collimator**
- **Technique**
- **Results**
- **Summary**



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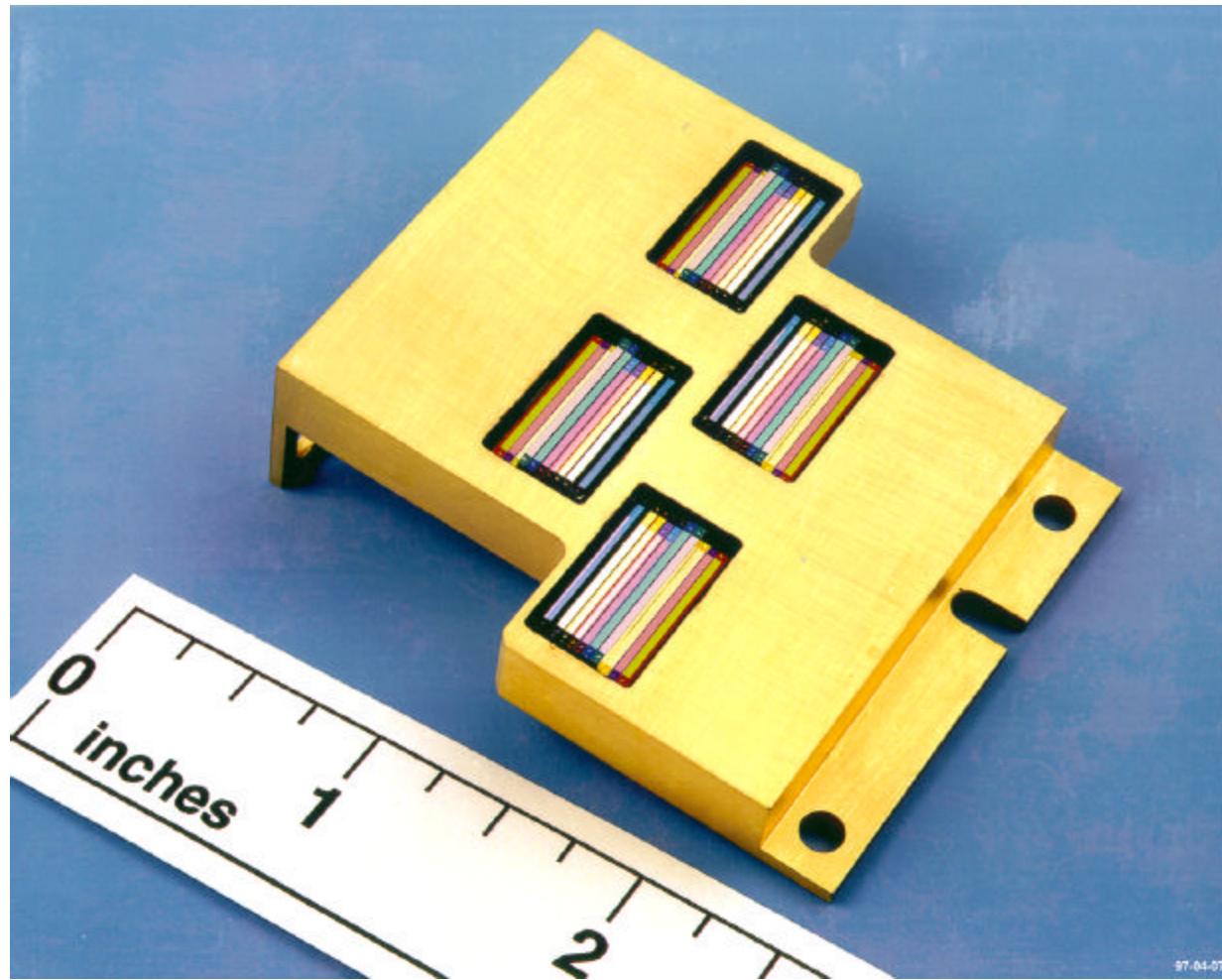
## EO-1 ALI MS/PAN Spectral and Spatial Coverage

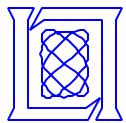
Band	Wavelength( <i>nm</i> )	Detector Type	GSD ( m )	# of Detectors
Pan	0.480-0.690	Si Photodiode	10	3840
MS-1'	0.433-0.453	Si Photodiode	30	1280
MS-1	0.450-0.515			
MS-2	0.525-0.605			
MS-3	0.630-0.690			
MS-4	0.775-0.805			
MS-4'	0.845-0.890			
MS-5'	1.200-1.300	PV HgCdTe	30	1280
MS-5	1.550-1.750			
MS-7	2.080-2.350			



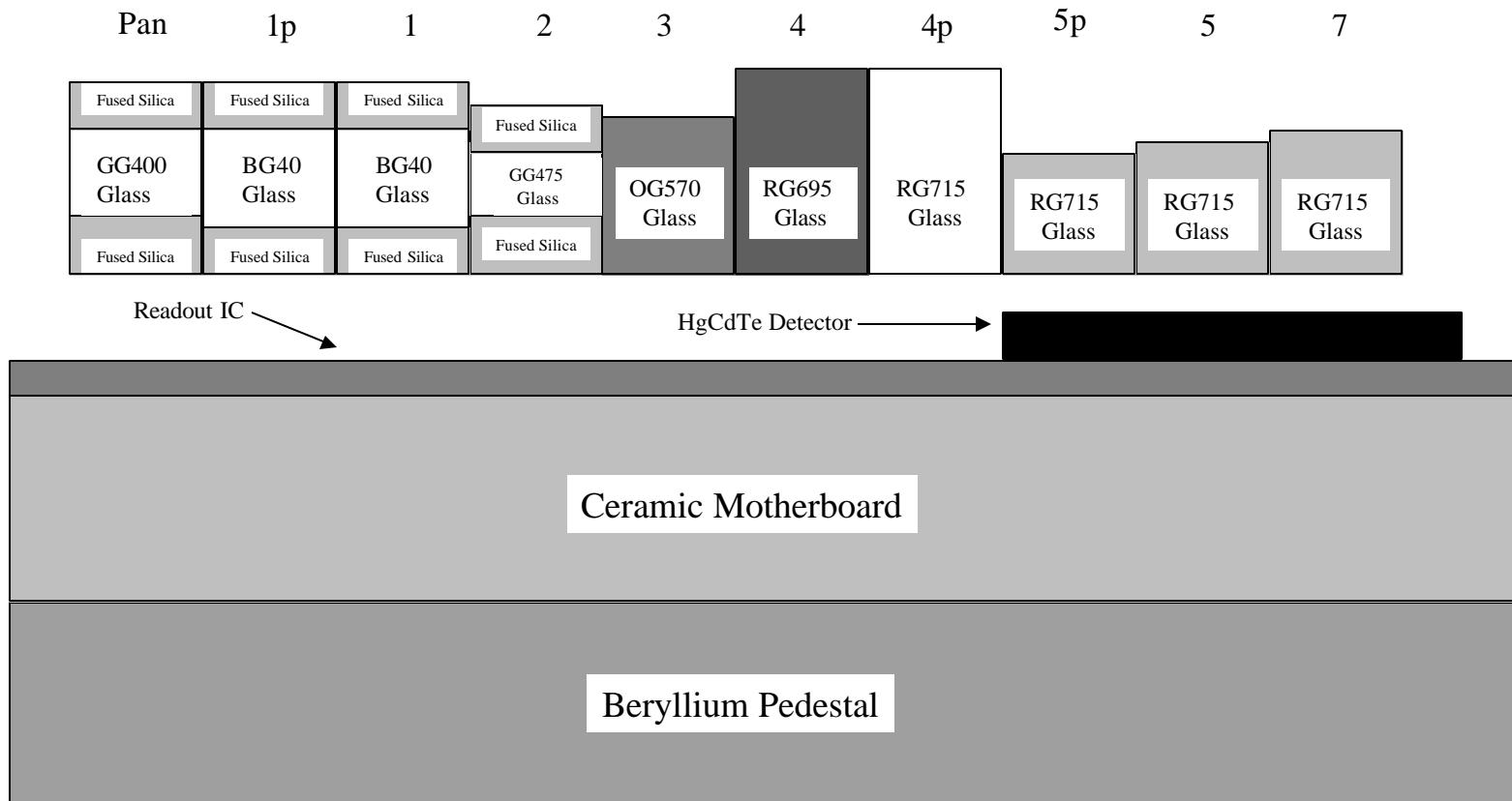
# MS/PAN Flight Module

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# Spectral Filters

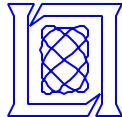




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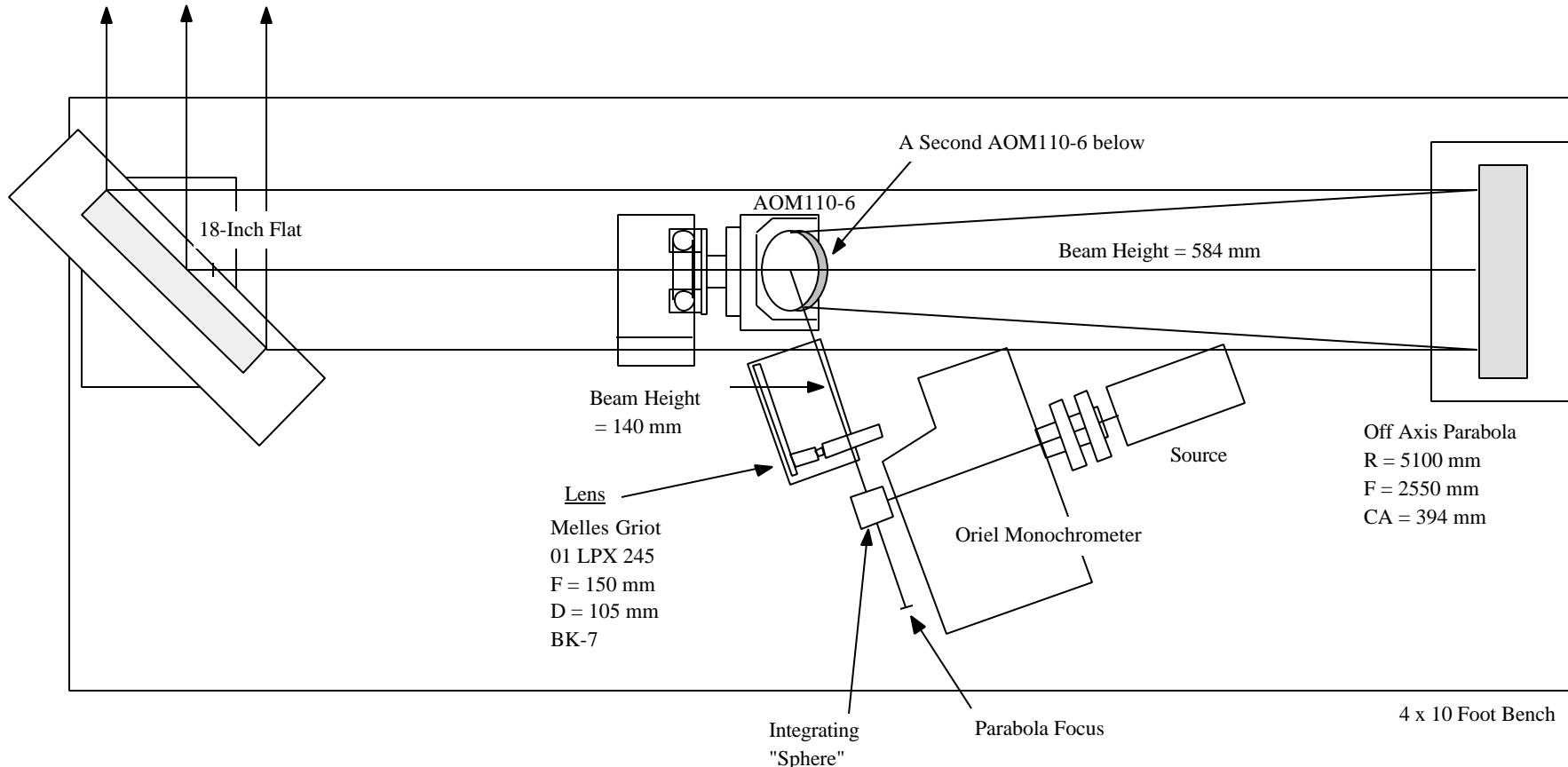
# Spectral Collimator

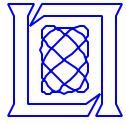
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- **Source**
  - Quartz tungsten halogen lamp
  - Oriel MS257 monochromator
  - **2” diameter (id) integrating sphere**
    - 0.5” diameter exit port
- **Collimating Optics(0.5° field)**
  - 6” diameter Condensing lens
  - 17” diameter, 100” focal length off-axis parabola
  - 20” diameter turning flat
- **Careful baffling to ensure proper stray light rejection**



# Spectral Collimator

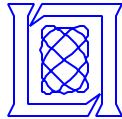




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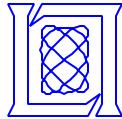
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# Spectral Calibration Technique

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- **Measure spectral response of multispectral and panchromatic detectors from 0.4-2.5 mm**
- **Technique**
  - Flood focal plane with 2 nm FWHM beam sequentially from 350 to 1000 nm in 2 nm intervals for VNIR
  - Flood focal plane with 4 nm FWHM beam sequentially from 1000 to 2500 nm in 4 nm intervals for SWIR
  - Record response of focal plane at each interval
  - Record response of spectrally calibrated silicon and lead sulfide detectors to monitor flux and stability of beam



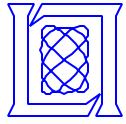
# Spectral Calibration Analysis

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- A pixel's spectral response is derived by normalizing its response as a function of wavelength to account for various artifacts
  - Dark current
  - Vacuum tank window transmission
  - Beam flux

$$S_p(b, l) = \frac{R_p(dn, l) R_d(l)}{T_w(l) F(l)}.$$

- $S_p(b, l)$  is the derived spectral response for pixel P as a function of band b and wavelength l
- $R_p(dn, l)$  is the ALI dark subtracted focal plane response for pixel P as a function of wavelength
- $R_d(l)$  is the spectral responsivity of the reference detector used to monitor the beam
- $T_w(l)$  is the spectral transmission of the vacuum tank window
- $F(l)$  is the measured flux of the beam as a function of wavelength



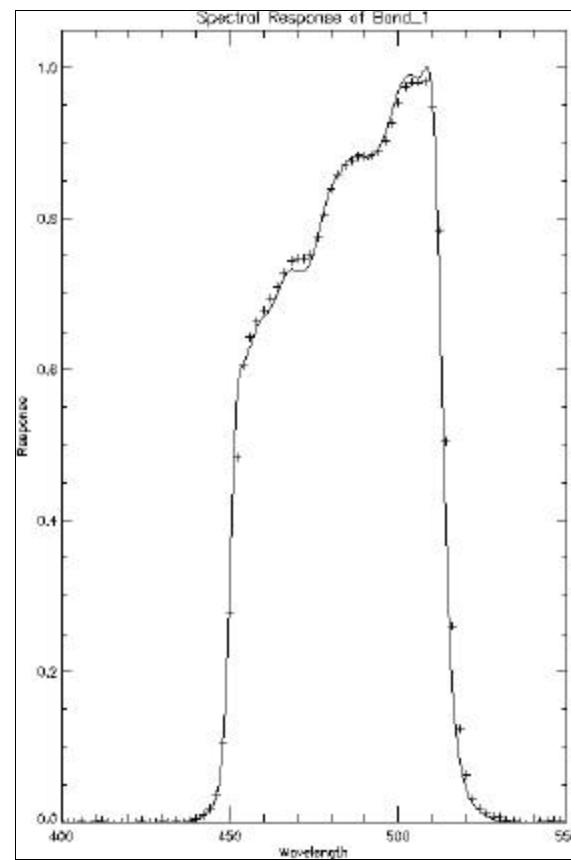
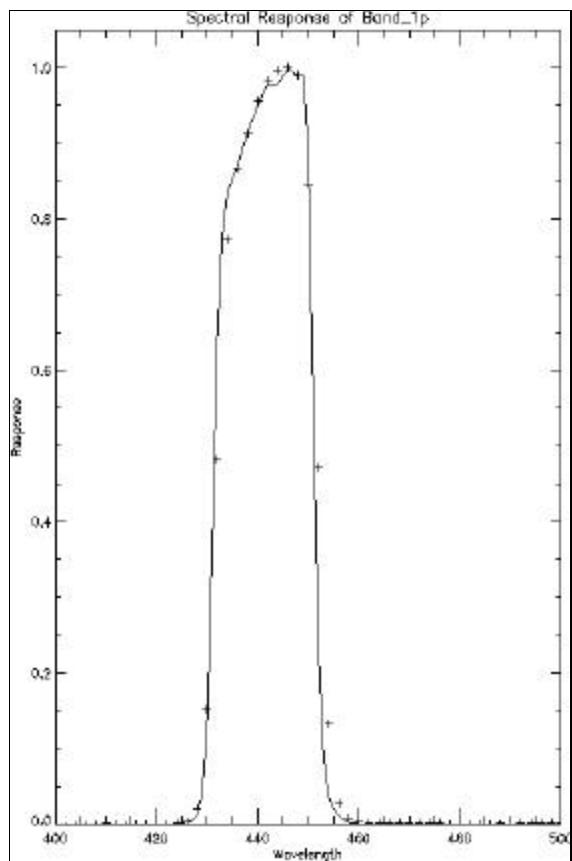
# Outline

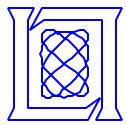
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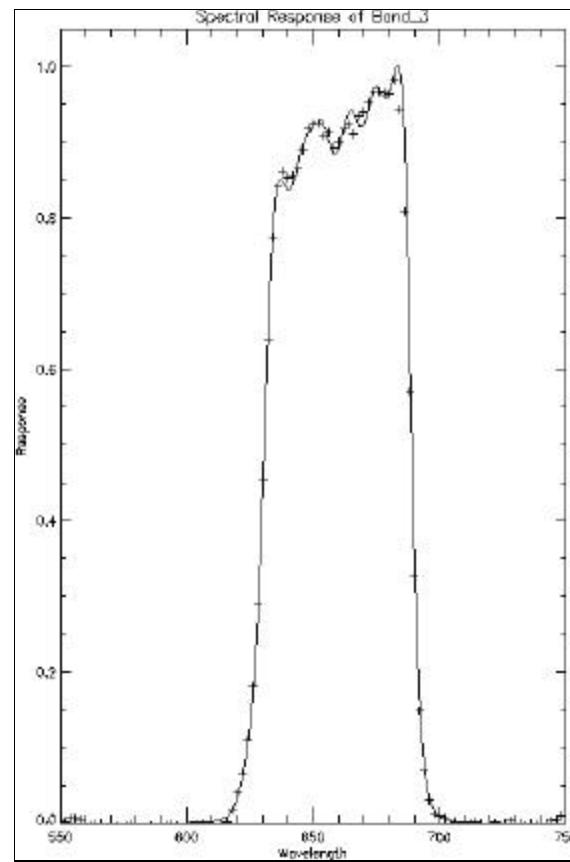
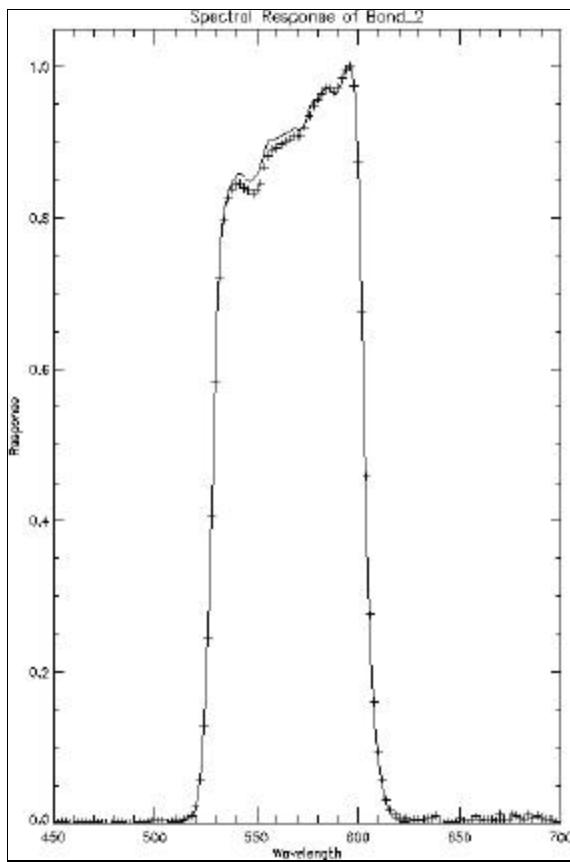
# Spectral Calibration Results: Band 1p and 1

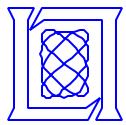




# Spectral Calibration Results: Bands 2 and 3

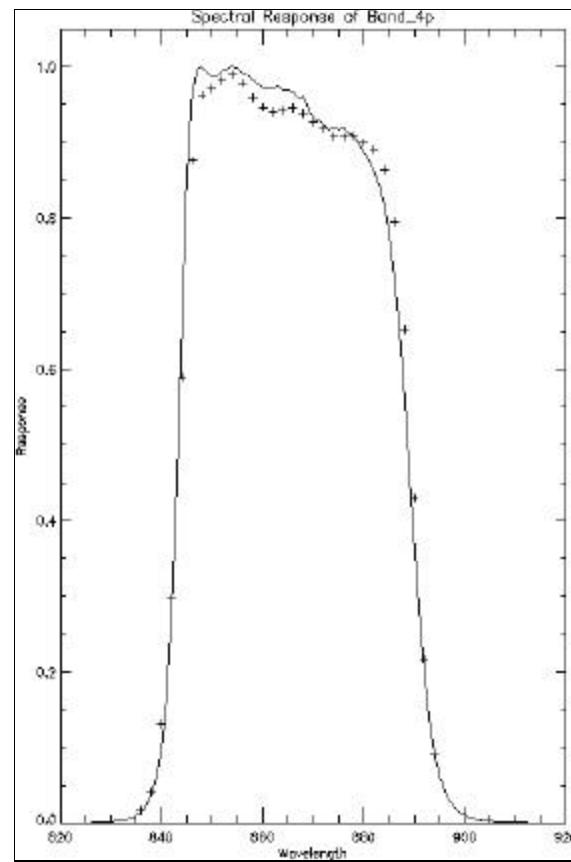
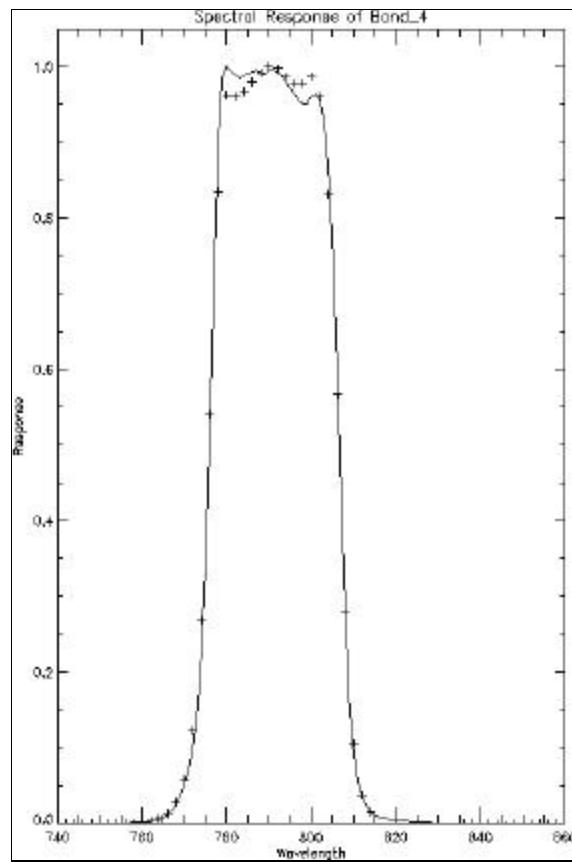
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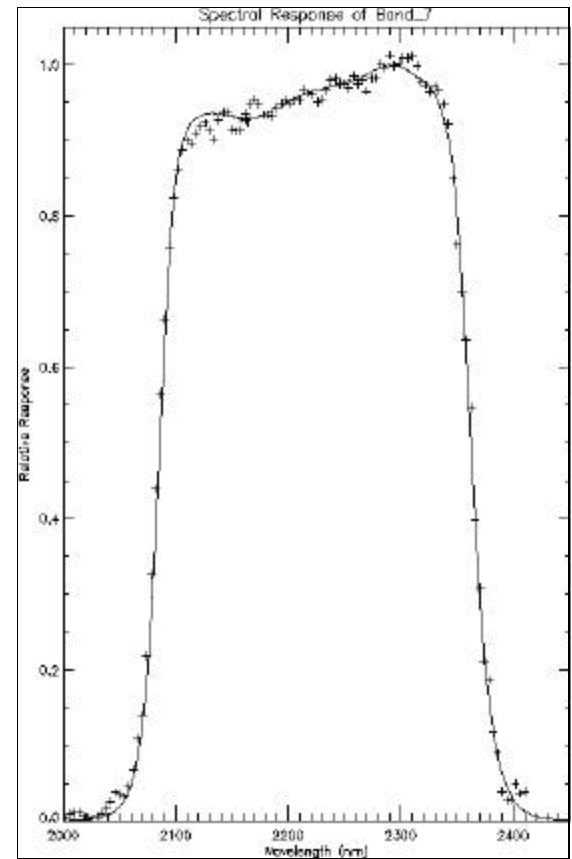
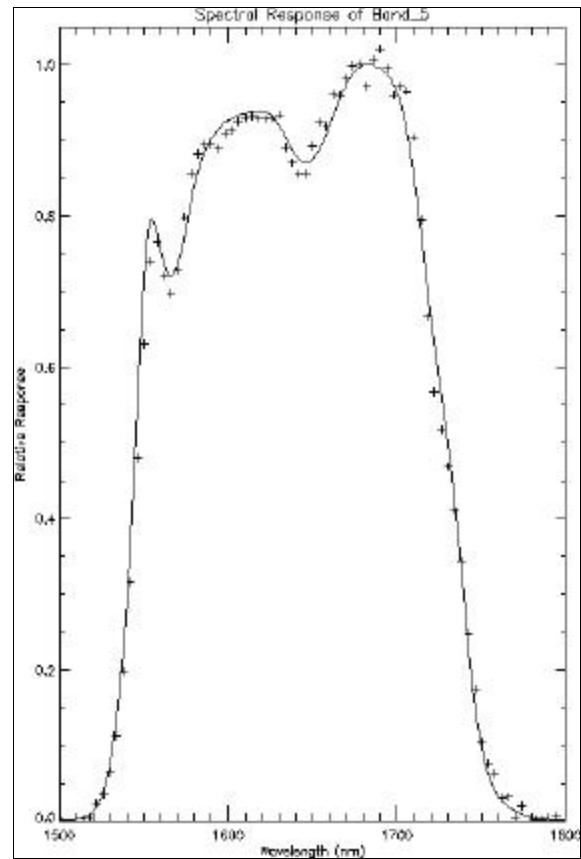
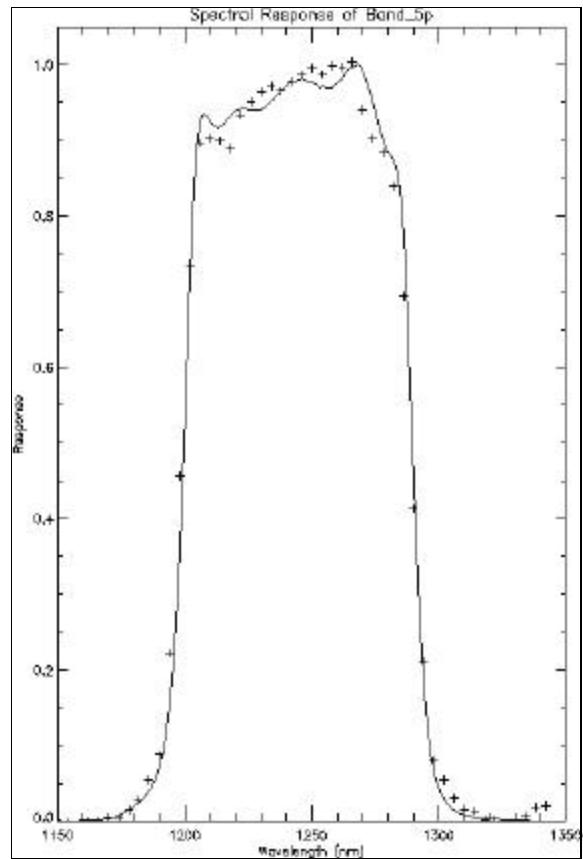
# Spectral Calibration Results: Bands 4 and 4p

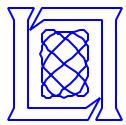
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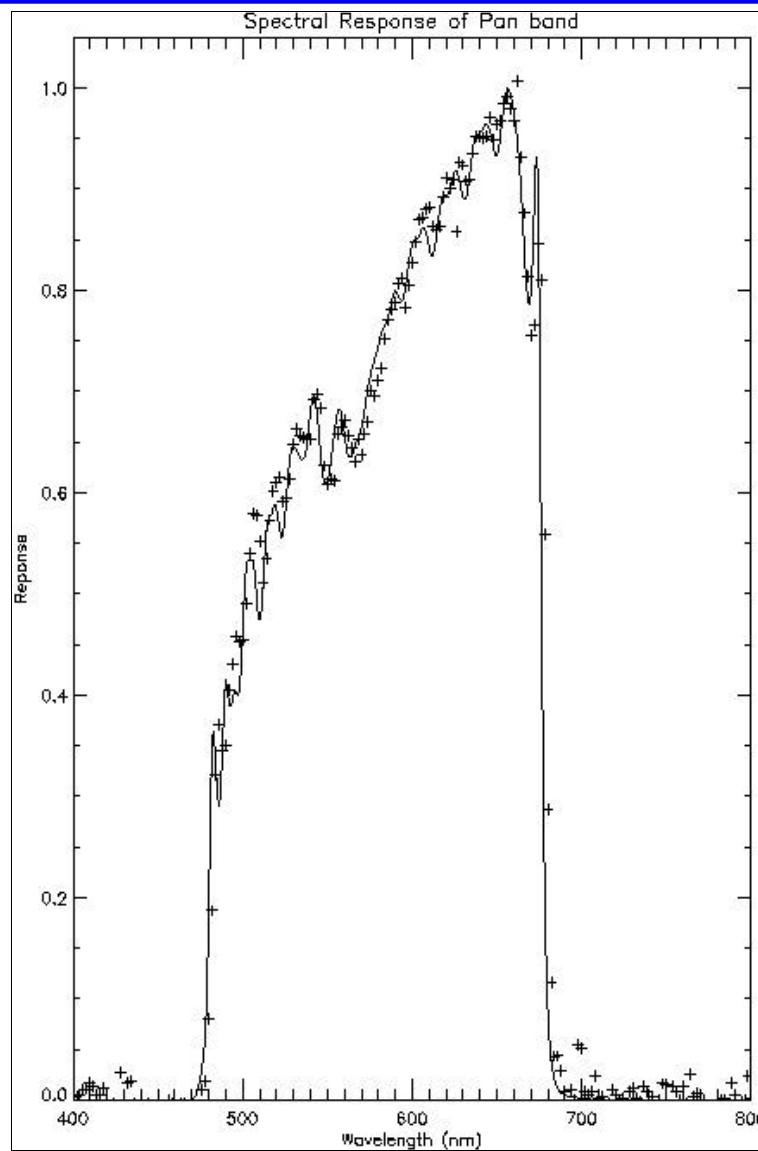


# Spectral Calibration Results: Bands 5p, 5 , and 7



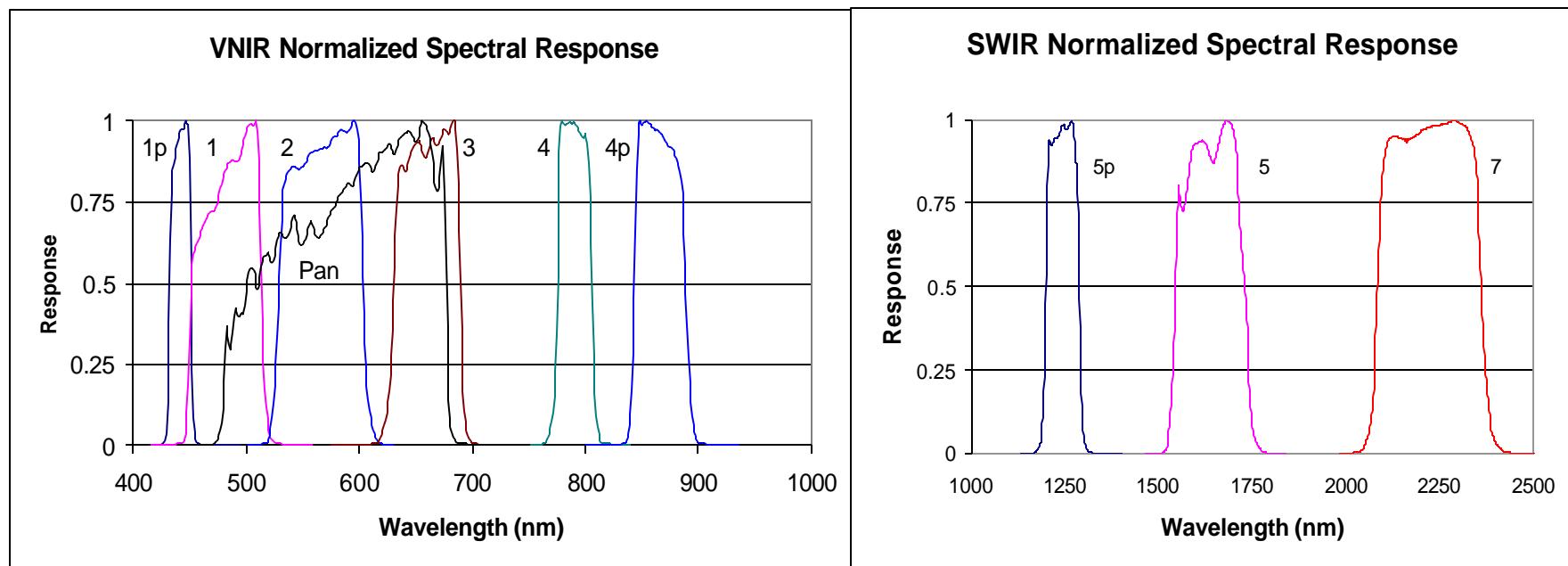


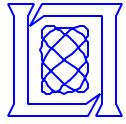
# Spectral Calibration Results: Panchromatic Band





# ALI Spectral Response Functions





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# Summary

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- We find excellent morphological agreement between system and subsystem level spectral calibration measurements.
- Based on system level measurements, the finer subsystem level calculations were adopted for the flight spectral response functions of the ALI.
- Characterization of a fully populated array would require substantial commitment at the system level and may be more appropriate at the subsystem level.